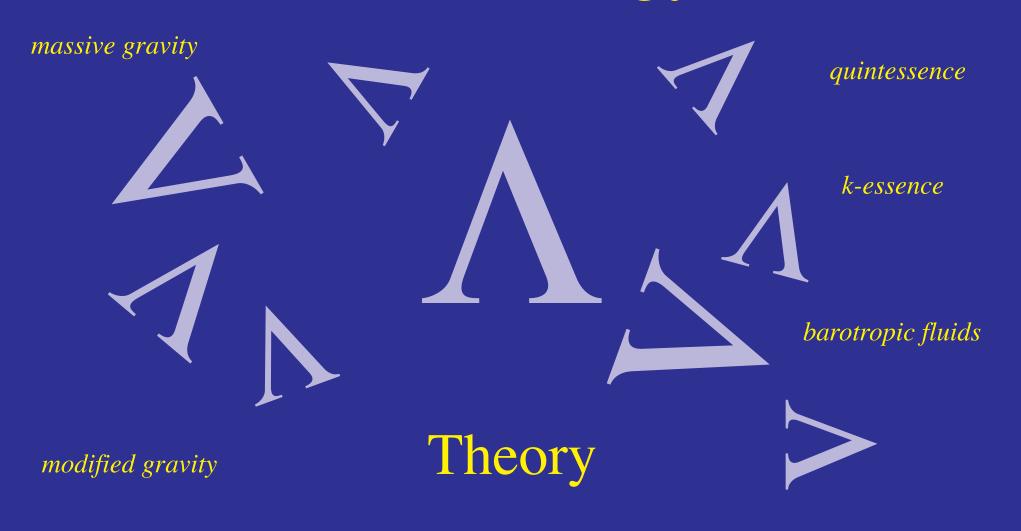
# Dark Energy



coupled quintessence

Wayne Hu SUSY, August 2011 frustrated defects

### Outline

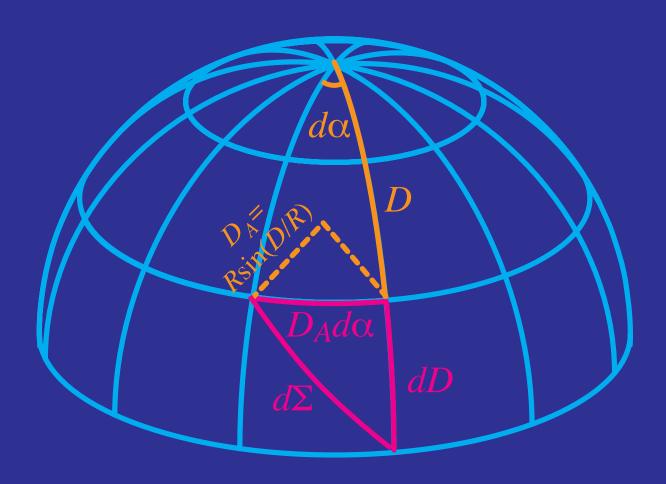
- Accelerating expansion history
- Dark energy phenomenology
- Cosmological constant
- Quintessence
- Couplings, modified forces
- Modified gravity

### Cosmic Acceleration

## Homogeneity & Isotropy

Homogeneity and isotropy → FRW line element

$$ds^{2} = -dt^{2} + a^{2}(t) \left[ dD^{2} + D_{A}^{2} d\alpha^{2} \right]$$
$$d\alpha^{2} = d\theta^{2} + \sin^{2}\theta d\phi^{2}$$



### Distance-Redshift

Photons travel on null-geodesics

$$D = \int \frac{dt}{a} = \int \frac{da}{aH} = \int \frac{dz}{H}$$

where the expansion rate  $H = \dot{a}/a$  and redshift  $(1+z) = a^{-1}$ 

• Given a measure of distance D (which measures time t) to an object at redshift z (which measures size a) infer the expansion history of universe a(t)



### Acceleration

• Relative distances between high and low z supernovae are further  $H_0D = \int dz (H_0/H)$  than expected in a matter only universe  $\rightarrow$  expansion rate

does not increase with redshift
does not decrease with expansion *a*as quickly

• If expansion rate drops with a as  $H \propto a^{-3(1+w)/2}$  then

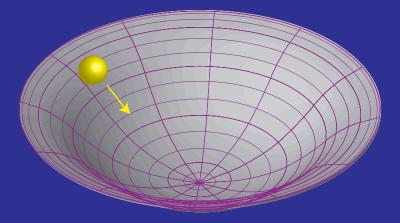
$$\frac{\ddot{a}}{a} \propto -(1+3w)a^{-3(1+w)}$$

acceleration if w < -1/3 or expansion rate falls slower than  $H \propto a^{-1}$ 

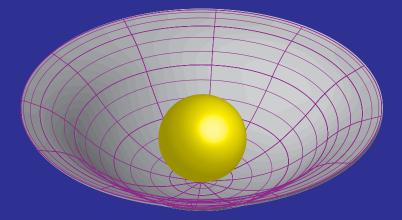
• Purely geometric so far, inference about missing dark energy requires Friedmann equation (Einstein equations)

### Mercury or Pluto?

General relativity says Gravity = Geometry



And Geometry = Matter-Energy



• Could the missing energy required by acceleration be an incomplete description of how matter determines geometry?

# Dark Energy

### Dark Energy

Under the Einstein equations

$$G_{\mu\nu} = 8\pi G T_{\mu\nu}$$

and the FRW metric, 00 and ii give the Friedmann equations

$$H^{2} = \frac{8\pi G}{3}(\bar{\rho} + \rho_{K})$$
$$\frac{\ddot{a}}{a} = -\frac{4\pi G}{3}(\bar{\rho} + 3\bar{p})$$

where  $\bar{\rho}$  and  $\bar{p}$  are the average energy density and pressure and are the only things allowed in  $T_{\mu\nu}$  by symmetry

- The index  $w = \bar{p}/\bar{\rho}$  is known as the equation of state parameter
- Observationally  $w \approx -1$  to  $\sim 10\%$  ruling out some possibilities like frustrated domain walls and strings

### $w \neq \text{Equation of State}$

- $w = \bar{p}/\bar{\rho}$  is a relationship between average energy density and pressure not local energy density and pressure
- Consider a barotropic fluid  $p(\rho)$

$$\bar{p} = w\bar{\rho} \to \nabla p = w\nabla \rho$$

which for 0 < w < -1 is an imaginary sound speed and violently unstable for constant w

 The only baryotropic equation of state that is allowed and accelerates the expansion is

$$p(\rho) = f(\rho - \text{const}) - \text{const}.$$

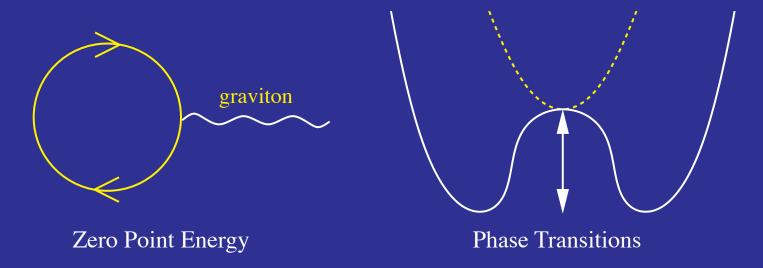
with f' > 0 – "Chaplygyn gas" models RIP

• Even more generally beyond  $p(\rho)$ : one non-dynamical dominant piece, zero or more subdominant dynamical piece

# Cosmological Constant

### Cosmological Constant

- Simplest possibility, consistent with all data to date, is a constant: Einstein's Cosmological Constant
- Particle physics provides sources for such a constant



- But the energy scales associated with particle physics scale cutoffs and transitions give energy densities ( $\rho \sim E^4$ ) at least  $\sim 60$  orders of magnitude too large
- For a bare CC to cancel these contributions would seem to require exquisite fine tuning

## String Landscape?

- String landscape provides ~10<sup>500</sup> metastable vacua [Kachru, Kallosh, Linde, and Trivedi 2003]
- At some of these, this cancellation is achieved [so if they can be populated and selected anthropically...]



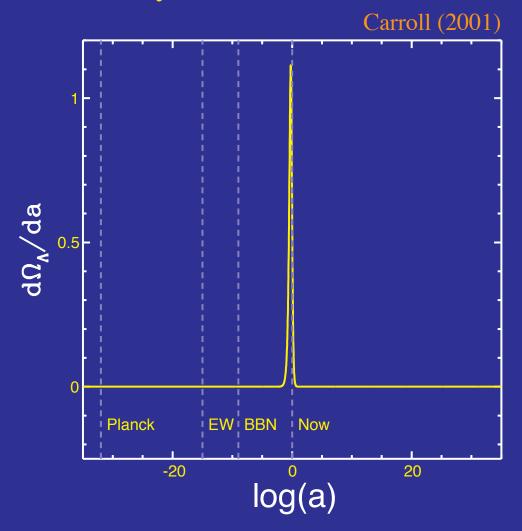
## String Landscape?

- String landscape provides ~10<sup>500</sup> metastable vacua [Kachru, Kallosh, Linde, and Trivedi 2003]
- Meets some resistance from people with flat backgrounds who think landscapes are artificial!



### Coincidence

Anthropic arguments attempt to address coincidence problem:
 matter/radiation dilutes with expansion
 dark energy constant or slowly diluting
 only comparable today



# Quintessence

### Quintessence

• Perhaps the true cosmological constant is zero and we are rolling in a (very!) flat direction of a landscape like inflation [but what protects a  $m\sim H_0\sim 10^{-33} \text{eV}$  mass and small couplings?]



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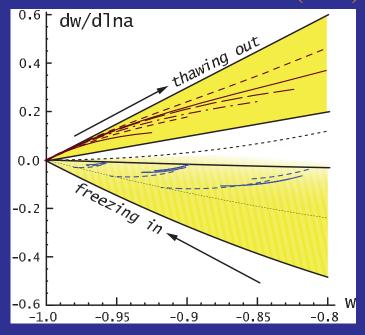
[possibly trading coincidence with features in potential]

- Two degrees of freedom:

   potential energy (driving acceleration)
   kinetic energy (associated with rolling)
   dynamical dark energy
- Typical models:

   thawing frozen by Hubble drag,
   released to roll
   freezing
   rolling/tracking early
   on and slowing to potential domination

#### Caldwell & Linder (2005)



# Beyond the Background

### Beyond w

- Hallmark of cosmological constant / quintessence spatial smoothness relative to matter
  - for quintessence, sound speed of kinetic contribution  $c_S$ = 1 smooth inside the horizon
- Beyond a minimally coupled light scalar field: self-interactions, coupling to dark matter or baryons etc.
- From the matter standpoint these look like changes to the (gravitational) force law
- Currently, strongest evidence for acceleration is from expansion history
- Tests of growth of structure are starting to be incisive and can confirm or refute predictions of the paradigm

### Post-Friedmann Parameterization

- Repeat the logic of the background geometry calculation
- Photons move on null geodesics of the perturbed post Friedmann metric

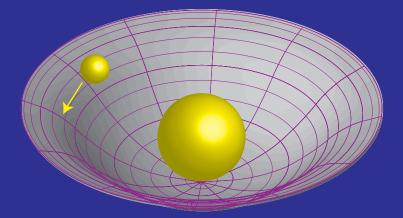
$$ds^{2} = -(1+2\Psi)dt^{2} + a^{2}(1+2\Phi)\left[dD^{2} + D_{A}^{2}d\alpha^{2}\right]$$

and feels both the Newtonian potential  $\Psi$  and space curvature potential  $\Phi$ 

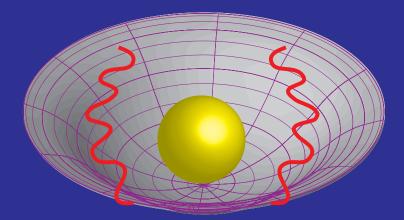
- Non-relativistic matter moves according to  $\Psi$  and sources metric fluctuations via Einstein equation (Poisson equation)
- Modified forces and couplings can break these relations and dark energy perturbations can provide additional sources

# Dynamical vs Lensing Mass

• Newtonian potential:  $\Psi = \delta g_{00}/2g_{00}$  which non-relativistic particles feel



• Space curvature:  $\Phi = \delta g_{ii}/2g_{ii}$  which also deflects photons



 Most of the incisive tests of gravity reduce to testing the space curvature per unit dynamical mass

### Modified Gravity = Dark Energy?

- Solar system tests of gravity are informed by our knowledge of the local stress energy content
- With no other constraint on the stress energy of dark energy other than conservation, modified gravity is formally equivalent to dark energy

$$F(g_{\mu\nu}) + G_{\mu\nu} = 8\pi G T_{\mu\nu}^{M} - F(g_{\mu\nu}) = 8\pi G T_{\mu\nu}^{DE}$$

$$G_{\mu\nu} = 8\pi G [T_{\mu\nu}^{M} + T_{\mu\nu}^{DE}]$$

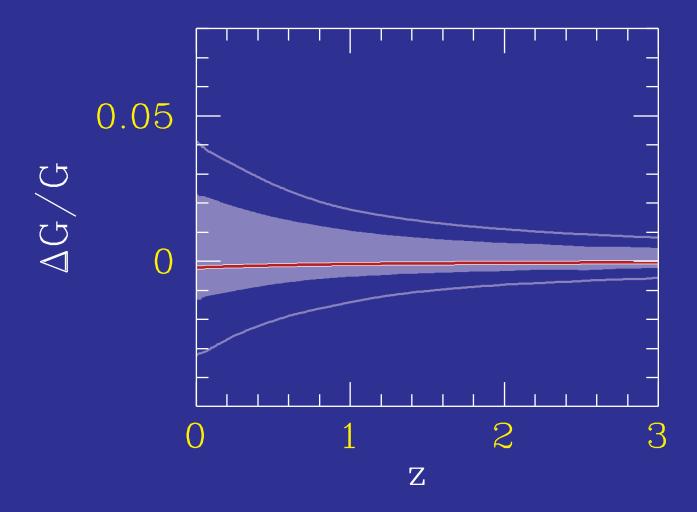
and the Bianchi identity guarantees  $\nabla^{\mu}T^{\rm DE}_{\mu\nu}=0$ 

- Distinguishing between dark energy and modified gravity requires closure relations that relate components of stress energy tensor
- For matter components, closure relations take the form of equations of state relating density, pressure and anisotropic stress

# Geometry Predicts Growth

# Falsifying ACDM

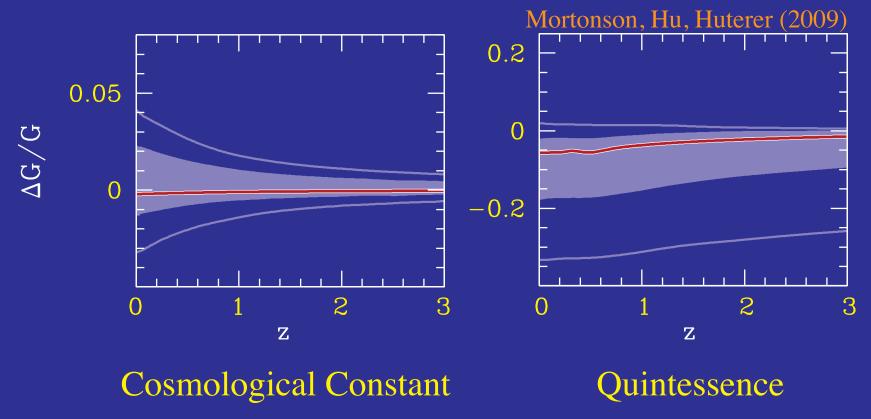
• A slows growth of structure in highly predictive way



Cosmological Constant

### Falsifying Quintessence

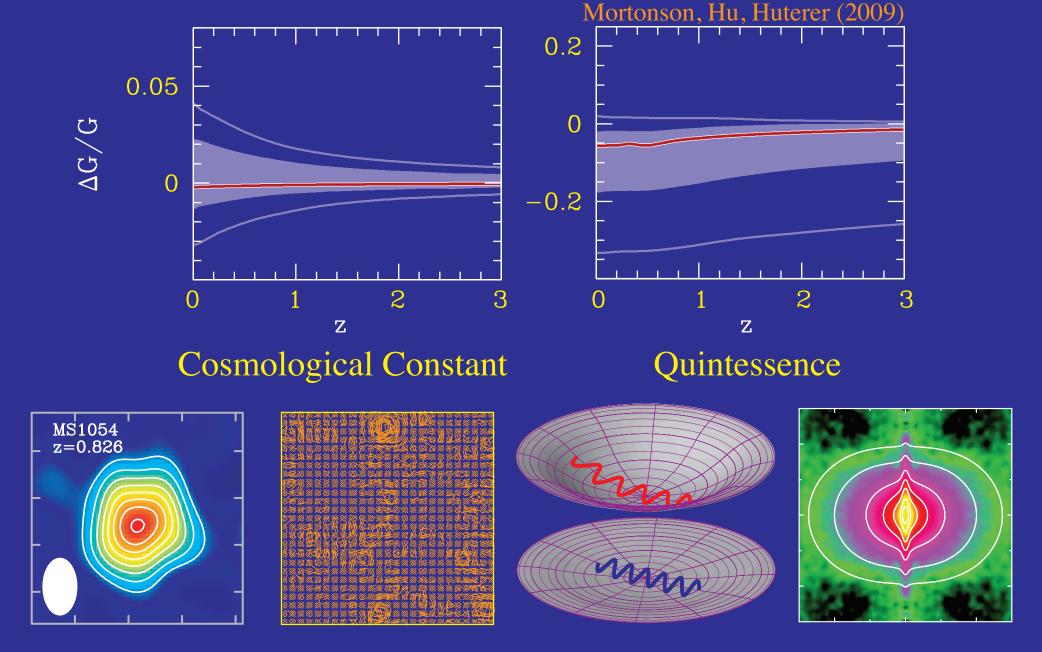
Dark energy slows growth of structure in highly predictive way



- Deviation significantly > 2% rules out  $\Lambda$  with or without curvature
- Excess >2% rules out quintessence with or without curvature and early dark energy [as does >2% excess in  $H_0$ ]

### Dynamical Tests of Acceleration

Dark energy slows growth of structure in highly predictive way



### Modified Forces

### Modified Forces

- Extra scalar propagating degree of freedom
- Cosmological IR modification hidden from local constraints on gravity and fifth forces → non-linear mechanism (strong interactions or changes in the potential or coupling)

Chameleon mechanism (running mass or coupling)
Vainshtein mechanism (strong coupling, derivative interactions)

• Concrete (but toy) models that exhibit these Modified Action f(R)

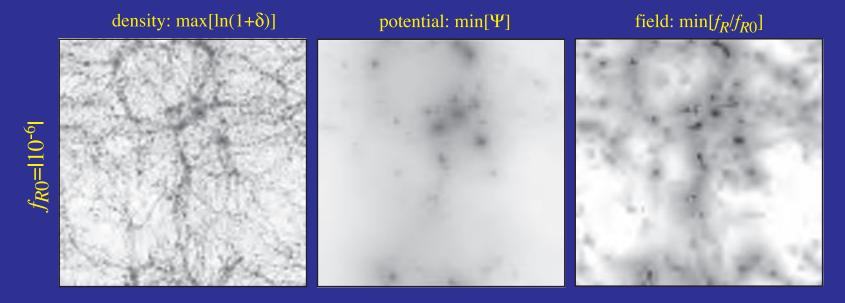
$$S = \int d^4x \sqrt{-g} \left[ \frac{R + f(R)}{16\pi G} + \mathcal{L}_{\rm m} \right]$$

Dvali-Gabadadze-Porrati (DPG) Braneworld

$$S = \int d^5x \sqrt{-g} \left[ \frac{^{(5)}R}{2\kappa^2} + \delta(\chi) \left( \frac{^{(4)}R}{2\mu^2} + \mathcal{L}_m \right) \right]$$

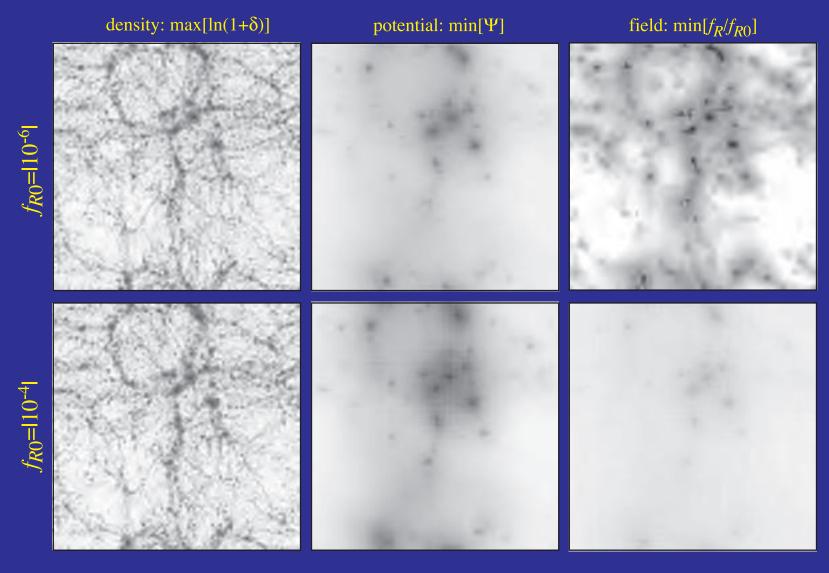
### Environment Dependent Force

• Chameleon suppresses extra force (scalar field) in high density, deep potential regions



### **Environment Dependent Force**

• For large background field, gradients in the scalar prevent the chameleon from appearing



### Massive Gravity

- DGP model motivated re-examination of massive gravity models [de Rham, Gabadadze, et al, Koyama et al, Hassan & Rosen (2010-2011)
- Graviton mass  $\sim H_0$  provides self-acceleration

$$H^2 = m^2 + \frac{8\pi G}{3}\rho$$

while also not seeing the cosmological constant contribution "degravitation"

• Key: add extra terms to Fierz-Pauli action to make it nonlinearly ghost free [Arkani-Hamed, Georgi, Schwartz (2003], exhibit Vainshtein strong coupling (Galileon symmetry, restoring vDVZ continuity)

$$S = \int d^4x \sqrt{-g} \frac{1}{16\pi G} \left( R + m^2 [\mathcal{L}^{(2)}(\mathcal{K}) + \alpha_3 \mathcal{L}^{(3)}(\mathcal{K}) + \alpha_4 \mathcal{L}^{(4)}(\mathcal{K})] \right)$$
with  $\mathcal{K}^{\mu}_{\nu} = \delta^{\mu}_{\nu} + \sqrt{g^{\mu\alpha} \partial_{\alpha} \phi^a \partial_{\nu} \phi^b \eta_{ab}}$ 

Much progress in the last year! stay tuned...

### Summary

- Strong evidence for cosmic acceleration from distance-redshift or geometric probes
- Einstein/Friedmann equations imply negative pressure component  $\bar{p}/\bar{\rho} < -1/3$
- Dark energy looks like const (+ small dynamical component) must look phenomenologically like cosmological constant
- Why small but finite cosmological constant? string landscape?
- Quintessence: dynamics of minimally coupled, very light, slowly rolling scalar field
- Couplings → fifth forces, difficult to hide Chameleon, Vainshtein
- Fifth forces subset of "modified gravity" explanations
- Recent progress on making massive gravity explain acceleration, cosmological constant problem...